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# The use of the correlation function technique for the Anderson transition 

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#### Abstract

For a completely disordered lattice, it is found that vanishing mobility is attained at a critical atomic density of $\rho_{\mathrm{A}}^{1 / 3} a_{0}=0.2$, which agrees adequately with known results.


## 1. Introduction

The Anderson transition is a property of disordered conductors whereby the system goes from a metallic to an insulating state as the disorder is increased. To study this phenomenon one starts with a model of the conducting state in which the disorder may be due either to lattice irregularity (Matsubara and Toyozawa 1961) or site energies (Edwards and Thouless 1972) or both (Mott and Davis 1971). One chooses a parameter to measure the degree of each type of irregularity and changes the parameters to vary the randomness of the system.

There are different methods to study the transition regime such as looking at the mean free path $L$ (Ioffe and Regel 1960, Debney 1976), using the $L(E)$ localization test (Economou and Cohen 1972), and using correlation function techniques which directly look at the transport coefficients (Vollhardt and Wolfle 1980, Gotze 1981). Calculating the transport coefficients from a suitable formula is a direct way to examine the above transition, since a vanishing value of the coefficient is indicative of the insulating state. Currently, new and practical formulae are being developed for transport coefficients (Milinski 1991, Unal et al 1992) and their capabilities should be tested through various applications.

In this work we apply the newly derived mobility formula of Unal et al (1992)

$$
\begin{equation*}
\mu=\frac{1}{2} e N \frac{\alpha_{r}^{\prime \prime}}{\alpha_{r}^{2}} \frac{1}{3} \tag{1.1}
\end{equation*}
$$

to a completely disordered lattice. Here $e$ is the electronic charge, $N$ is the total number of electrons and $\alpha_{r}{ }^{\prime}, \alpha_{r}{ }^{\prime \prime}$ are the frequency derivatives of the correlation function

$$
\begin{equation*}
\alpha(\omega)=\left\langle F_{x}\right| \frac{-j}{\omega-j s-L}\left|F_{x}\right\rangle=\alpha_{r}(\omega)+j \alpha_{i}(\omega) \tag{1.2}
\end{equation*}
$$

where

$$
\begin{align*}
& \alpha_{r}(\omega)=\pi\left\langle F_{x}\right| \delta(\omega-L)\left|F_{x}\right\rangle  \tag{1.3}\\
& \alpha_{i}(\omega)=-\left\langle F_{x}\right| \frac{1}{\omega-L}\left|F_{x}\right\rangle . \tag{1.4}
\end{align*}
$$

In the above equations, $L$ is the usual Liouville operator, $\omega$ represents frequency and $F_{x}$ is the internal force related to the momentum $P_{x}$ through $F_{x}=j L P_{x}$. It has been shown (Unal and Alkan 1993) that $\alpha_{r}(\omega)$ can be put into the form

$$
\begin{equation*}
\left.\alpha_{r}(\omega)=-\frac{2 \pi}{3} \sum_{k q} \hbar q^{2} \right\rvert\, U_{q}\left[^{2} n_{k}^{\prime} \delta\left(\hbar \omega-\Delta E^{\prime}\right)\right. \tag{1.5}
\end{equation*}
$$

The meaning of each term in this expression will be defined in the following sections.
The density of the medium, $\rho_{\mathrm{A}}$, is considered in this paper to be the randomness parameter, and as we increase it we observe that the mobility drops to zero, thus giving the onset of the Anderson transition. As will be discussed later, the critical atomic density we obtain, $\rho_{\mathrm{A}}^{1 / 3} a_{0}=0.2$, agrees reasonably with other estimates (Unal 1987, Debney 1976). We take this to be an indication that (1.1) works very well in this case.

Section 2 sets out the model Hamiltonian, section 3 is about the evaluation of $\alpha_{r}{ }^{\prime}, \alpha_{r}{ }^{\prime \prime}$ and section 4 deals with the transition density through an examination of $\mu$. Finally, in the appendix we present detailed calculation of some integrals.

## 2. The model Hamiltonian

In this paper we consider the scattering of electrons in a metallic substance whose lattice site positions $\boldsymbol{R}_{i}$ have no regularity, which is known as a completely disordered lattice. Since this model is described fully elsewhere (Matsubara and Toyozawa 1961) we shall present here only an outline. Each atom is assumed to have an effective Bohr radius $a_{0}$ and is supposed to give a is electron to the medium. The mean atomic distance $a$ is given by $a^{3}=\Omega / N=\rho^{-1}$ where $\Omega$ and $N$ show the volume and the number of atoms, respectively, and where $\rho$ is the atomic density. In the quasi-momentum representation the system Hamiltonian $H=H_{0}+U$ has the forms

$$
\begin{align*}
& H_{0}=\sum_{k} H_{k k} C_{k}^{+} C_{k}  \tag{2.1}\\
& U=\sum_{k q} H_{k+q, k} C_{k+q}^{+} C_{k} \tag{2.2}
\end{align*}
$$

where $H_{k, k}=\left\langle k^{\prime}\right| H|k\rangle, k^{\prime}=k+q$, and $C_{k}^{+}, C_{k}$ are creation and annihilation operators for the electrons. Using the same ideas as in a previous work (Unal 1987) we can replace the diagonal elements of $H$ by their averages $\left\langle H_{k k}\right\rangle=(N / \Omega) V(k)=-\epsilon(k)$, where $V(k)$ is the Fourier transform of the interatomic hopping potentials and $\epsilon(k)$ is given by

$$
\begin{equation*}
\epsilon(k)=\frac{64\left(a_{0} / a\right) \pi}{\left(1+a_{0}^{2} k^{2}\right)^{3}} \text { Ryd. } \tag{2.3}
\end{equation*}
$$

Later in our calculation we shall need $\left.\left.\langle | H_{k^{\prime} k}\right|^{2}\right\rangle=(1 / N)\left|U_{q}\right|^{2}$, where $\left|U_{q}\right|^{2}$ is given by

$$
\begin{equation*}
\left|U_{q}\right|^{2}=2 \epsilon_{k}^{2}+2 \epsilon_{k}^{2}+\frac{7}{16} \mathcal{A}^{3}+\mathcal{A}^{3}\left(7+a_{0}^{2}\left|k+k^{\prime}\right|^{2}\right) / 16\left(1+a_{0}^{2}\left|k+k^{\prime}\right|^{2}\right)^{4} . \tag{2.4}
\end{equation*}
$$

Here $\mathcal{A}=4\left(a_{0} / a\right) \pi^{1 / 3}$ has been defined. The mobility equation contains the terms $\alpha_{r}{ }^{\prime}$ and $\alpha_{r}{ }^{\prime \prime}$, and we can write out these expressions, in our case, by comparing with similar
expressions in the work of Unal and Alkan (1993). Looking at their (2.6) and (2.7) we can write our $\alpha_{r}{ }^{\prime}, \alpha_{r}{ }^{\prime \prime}$ respectively as

$$
\begin{align*}
& \alpha_{r}^{\prime}=\frac{2 \pi}{3} \sum_{k q} \hbar^{2} q^{2}\left|U_{q}\right|^{2} n_{k}^{\prime} \delta^{\prime}\left(\Delta E^{\prime}\right)  \tag{2.5}\\
& \alpha_{r}^{\prime \prime}=-\frac{2 \pi}{3} \sum_{k q} \hbar^{3} q^{2}\left|U_{q}\right|^{2} n_{k}^{\prime} \delta^{\prime \prime}\left(\Delta E^{\prime}\right) \tag{2.6}
\end{align*}
$$

where $\Delta E^{\prime}=E_{k+q}-E_{k}, n_{k}$ is the occupation number and $n_{k}{ }^{\prime}$ stands for the derivative of the occupation number with respect to $E_{k}$. For reference purposes it has been necessary to use $k$ vectors multiplied by $\hbar$, but in fact units where $\hbar=1$ have been employed in this paper.

## 3. Evaluation of $\alpha_{r}{ }^{\prime}$ and $\alpha_{r}{ }^{\prime \prime}$

The sums in (2.5) and (2.6) may be evaluated by turning them into integrations in the usual way, by using the relations

$$
\sum_{q} \rightarrow \frac{V}{(2 \pi)^{3}} \int \mathrm{~d}^{3} q \quad \sum_{k} \rightarrow \frac{V}{(2 \pi)^{3}} \int \mathrm{~d}^{3} k
$$

If we denote the angle between the vectors $k$ and $q$ by $\alpha$, with $\Theta=\pi-\alpha, \cos \Theta=\nu$, the energy difference $\Delta E^{\prime}$ becomes

$$
\begin{equation*}
\Delta E^{\prime}=-\frac{\hbar^{2} k q}{m} \nu+\frac{\hbar^{2} q^{2}}{2 m} \tag{3.1}
\end{equation*}
$$

Integration is carried out in the order $\mathrm{d} \nu \rightarrow \mathrm{d} q \rightarrow \mathrm{~d} k$ because, with this order, during the transformations of delta functions into simpler forms, Jacobians never become zero; otherwise, nothing fixes this order. The evaluation of $\alpha_{r}{ }^{\prime}$ and $\alpha_{r}{ }^{\prime \prime}$ are given in the appendix, and the results are

$$
\left.\begin{array}{rl}
\alpha_{r}^{\prime}= & C_{\mathrm{I}}\{(1+
\end{array} \pi^{2} x^{2}\right)\left(1+4 \pi^{2} x^{2}\right)^{3}\left[16+\frac{7}{128}(x / \pi)\left(1+\pi^{2} x^{2}\right)^{6}\right)
$$

Here $C_{1}, C_{2}$ are constants and $x$ is a parameter defined as $x=a_{0} / a$. Using these expressions in (1.1) we may obtain the mobility, $\mu$, as a function of $x$. When the mean interatomic distance $a$ is large, corresponding to small values of $x$, the atoms in the system are far apart from each other and in this case we expect that the scattering of an electron should be weak. The mobility will then have a large value, but as the parameter $x$ is increased the scattering of an electron with atoms will take place more effectively; therefore we may expect a decrease in the mobility. In order to check this expectation we looked at $\alpha_{r}{ }^{\prime \prime}$ appearing in the numerator of the mobility equation (1.1): $\alpha_{r}^{\prime \prime}$ had a value of $384 C_{2}$ for $x=0$ and it decreased as we slowly increased the parameter $x$. The value of $x=0.2$ was just enough to make $\alpha_{r}{ }^{\prime \prime}$ (and hence $\mu$ ) zero. While doing these processes we observed that $\alpha_{r}^{\prime}$ was never zero. We thus concluded that the entire system changed its state from a conducting to a non-conducting one: an Anderson transition took place.

## 4. Conclusions

As we mentioned in the introduction, there is currently a great deal of effort in developing new and practical formulae for the calculation of transport coefficients of condensed matter. These formulae differ from each other by the ease with which an actual computation can be carried out in particular applications. The correlation function technique of Vollhardt and Wolfle (1980) and of Gotze (1981) was used by one of us in a previous work and gave the result $\rho_{\mathrm{A}}^{1 / 3} a_{0}=x=0.29$ (Unal 1987), but in this paper we use the mobility formula developed by Milinski (1991) and the corrected version developed by Unal et al (1992). The mobility formula has not been used before in the study of the Anderson transition for a completely disordered lattice model. In this respect our work aims to show the usefulness of the mobility formula for studying the Anderson transition. Using the Ioffe-Regel condition for the mean free path of electrons, Debney (1976) found that this transition took place at $x=0.33$ while Fertis et al (1981) and Logan and Volynes (1985) obtained the values $x=0.16$ and $x=0.12$ respectively. Our result in this paper, $x=0.20$, falls right inside the region subtended by the above values. This shows that the use of the mobility formula may constitute a basis for the study of the Anderson transition.

## Appendix.

In order to evaluate $\alpha_{r}^{\prime}$ the expressions of $\left|U_{q}\right|^{2}$ and $\Delta E^{\prime}$, given by (2.4) and (3.1) respectively, are used in (2.5). For the $n_{k}{ }^{\prime}$ term appearing in (2.5) we use the result

$$
\begin{equation*}
n_{k}^{\prime}=\left[1-\left(\frac{\mathrm{d} A}{\mathrm{~d} \epsilon_{k}}\right)_{\epsilon_{\mathrm{F}}}\right]^{-1} \delta\left(\epsilon_{\mathrm{F}}-\epsilon_{k}\right) \tag{A.1}
\end{equation*}
$$

obtained from (25) and (27a) of a previous work (Unal 1987) under the same approximations cited therein. Now $\alpha_{r}{ }^{\prime}$ becomes

$$
\begin{align*}
\alpha_{r}^{\prime}=-\frac{2 \pi}{3} \sum_{k q} & \hbar^{2} q^{2} \frac{1}{N}\left(2 \epsilon_{k}^{2}+2 \epsilon_{k}^{2}+\frac{7}{16} \mathcal{A}^{3}\right. \\
& \left.+\mathcal{A}^{3} \frac{7+a_{0}^{2}\left(4 k^{2}-4 k q \nu+q^{2}\right)}{16\left[1+a_{0}^{2}\left(4 k^{2}-4 k q \nu+q^{2}\right)\right]^{4}}\right) \frac{m^{2}}{\hbar^{2} k^{2} \hbar^{2} q^{2}} \delta^{\prime}\left(\nu-v_{0}\right) \frac{\delta\left(\epsilon_{\mathrm{F}}-\epsilon_{k}\right)}{1-A^{\prime}\left(\epsilon_{\mathrm{F}}\right)} \tag{A.2}
\end{align*}
$$

The term $\delta^{\prime}\left(\Delta E^{\prime}\right)$ has been transformed as follows (Unal 1987):

$$
\begin{equation*}
\delta^{\prime}\left(-\frac{\hbar k \cdot \hbar q}{m} \nu+\frac{\hbar^{2} q^{2}}{2 m}\right)=-\frac{m^{2}}{\hbar^{2} k^{2} \hbar^{2} q^{2}} \delta^{\prime}\left(\nu-\nu_{0}\right) \tag{A.3}
\end{equation*}
$$

and since the root $\nu_{0}=\hbar q / 2 \hbar k$ has to remain between 0 and 1 we have the limitation $\hbar q=2 \hbar k$ during the $\mathrm{d} q$ integral. A factor of $2 \pi$ comes from the $\mathrm{d} \phi$ integral and thus we have
$\alpha_{r}^{\prime}=-\frac{2 \pi}{3 N} \frac{2 \pi V m^{2}}{\hbar^{3}(2 \pi)^{3}} \sum_{k} \frac{1}{\hbar^{2} k^{2}} \frac{\delta\left(\epsilon_{\mathrm{F}}-\epsilon_{k}\right)}{1-A^{\prime}} \int_{0}^{2 \hbar k} \hbar^{2} q^{2} \mathrm{~d}(\hbar q) \int_{0}^{1} \delta^{\prime}\left(v-v_{0}\right)[] .\mathrm{d} \nu$.

The notation [.] denotes all the contents of the square brackets in (A.2). Equation (A.4) becomes

$$
\begin{align*}
\alpha_{r}^{\prime}=-\frac{V m^{2}}{3 N \hbar^{3} \pi} & \sum_{k} \frac{1}{\hbar^{2} k^{2}} \frac{\delta\left(\epsilon_{\mathrm{F}}-\epsilon_{k}\right)}{1-A^{\prime}} \int_{0}^{2 \hbar k} \hbar^{2} q^{2} \mathrm{~d}(\hbar q)\left[\left(\frac{\left(64 \pi a_{0} / a\right)^{2}}{\left(1+a_{0}^{2} k^{2}\right)^{6}}+14 \pi\left(a_{0} / a\right)^{3}\right)\right. \\
& \times \delta(1-\hbar q / 2 \hbar k)+\int_{0}^{1} \frac{\left(64 \pi a_{0} / a\right)^{2}}{\left[1+a_{0}^{2}\left(k^{2}-2 k q v+q^{2}\right)\right]^{6}} \delta^{\prime}\left(v-v_{0}\right) \mathrm{d} v \\
& \left.+2 \pi\left(a_{0} / a\right)^{3} \int_{0}^{1} \frac{7+a_{0}^{2}\left(4 k^{2}-4 k q v+q^{2}\right)}{\left[1+a_{0}^{2}\left(4 k^{2}-4 k q v+q^{2}\right)\right]^{4}} \delta^{\prime}\left(v-v_{0}\right) \mathrm{d} v\right] . \tag{A.5}
\end{align*}
$$

Integration of the second and third terms inside the large square brackets is done by observing the rule

$$
\begin{align*}
& f(x) \delta^{\prime}(x)=f\left(x_{0}\right) \delta^{\prime}(x)-f^{\prime}\left(x_{0}\right) \delta(x)  \tag{A.6}\\
& \alpha_{r}^{\prime}=-\frac{V m^{2}}{3 N \hbar^{3} \pi} \sum_{k} \frac{1}{\hbar^{2} k^{2}} \frac{\delta\left(\epsilon_{\mathrm{F}}-\epsilon_{k}\right)}{1-A^{\prime}}\left[\int_{0}^{2 \hbar k} \hbar^{2} q^{2} \mathrm{~d}(\hbar q)\left(\frac{\left(64 \pi a_{0} / a\right)^{2}}{\left(1+a_{0}^{2} k^{2}\right)^{6}}+14 \pi\left(a_{0} / a\right)^{3}\right)\right. \\
& \times 2 \hbar k \delta(\hbar q-2 \hbar k)+\frac{\left(64 \pi a_{0} / a\right)^{2}}{\left(1+a_{0}^{2} k^{2}\right)^{6}} 2 \hbar k \int_{0}^{2 \hbar k} \hbar^{2} q^{2} \mathrm{~d}(\hbar q) \delta(\hbar q-2 \bar{k}) \\
&-12 a_{0}^{2} k \frac{\left(64 \pi a_{0} / a\right)^{2}}{\left(1+a_{0}^{2} k^{2}\right)^{7}} \int_{0}^{2 \hbar k} \hbar^{3} q^{3} \mathrm{~d}(\hbar q) \\
&+2 \pi\left(a_{0} / a\right)^{3} \int_{0}^{2 \hbar k} \frac{7+a_{0}^{2}\left(4 k^{2}-q^{2}\right)}{\left[1+a_{0}^{2}\left(4 k^{2}-q^{2}\right)\right]^{4}} \hbar^{2} q^{2} 2 \hbar k \delta(\hbar q-2 \hbar k) \mathrm{d}(\hbar q) \\
&\left.-24 \pi\left(a_{0} / a\right)^{3} a_{0}^{2} k \int_{0}^{2 \hbar k} \frac{9+a_{0}^{2}\left(4 k^{2}-q^{2}\right)}{\left[1+a_{0}^{2}\left(4 k^{2}-q^{2}\right)\right]^{5}} \hbar^{3} q^{3} \mathrm{~d}(\hbar q)\right] . \tag{A.7}
\end{align*}
$$

The $\mathrm{d}(\hbar q)$ integrals are easily carried out, and while doing $V /(2 \pi)^{3} \int \mathrm{~d}^{3} k$ integration we have to take into account the relation

$$
\begin{equation*}
\delta\left(\epsilon_{\mathrm{F}}-\epsilon_{k}\right)=\frac{\left(1+a_{0}^{2} k_{\mathrm{F}}^{2}\right)^{4}}{384 \pi\left(a_{0} / a\right) a_{0}^{2} k_{\mathrm{F}}}\left[\delta\left(\hbar k-k_{\mathrm{F}}\right)+\delta\left(\hbar k+k_{\mathrm{F}}\right)\right] . \tag{A.8}
\end{equation*}
$$

After the $\mathrm{d} k$ integration we have to put $k=k_{\mathrm{F}}$, and on putting $a=\pi / k_{\mathrm{F}}$ we can arrange the whole expression in terms of the parameter $x=a_{0} / a$. The result is

$$
\begin{align*}
\alpha_{r}^{\prime}=\frac{V m^{2}}{3 N \hbar^{3} \pi^{2}} & \frac{4 \pi V}{(2 \pi)^{3} \hbar^{3}} \frac{\left(1+\pi^{2} x^{2}\right)^{4}}{384\left(1-A^{\prime}\right) a_{0}^{3}} \llbracket 16(64 \pi x)^{2}\left(1+\pi^{2} x^{2}\right) \pi^{2} x^{2}\left(1+4 \pi^{2} x^{2}\right)^{3} \\
& +224 \pi^{3} x^{5}\left(1+4 \pi^{2} x^{2}\right)^{3}\left(1+\pi^{2} x^{2}\right)^{7}-48 \pi^{4} x^{4}(64 \pi x)^{2}\left(1+4 \pi^{2} x^{2}\right)^{3} \\
& -\pi x^{3}\left\{\left(9+4 \pi^{2} x^{2}\right)\left[3\left(1+4 \pi^{2} x^{2}\right)^{4}-4\left(1+4 \pi^{2} x^{2}\right)^{3}+1\right]\right. \\
& -3\left(1+4 \pi^{2} x^{2}\right)^{5}+\pi^{2} x^{2}\left(1+4 \pi^{2} x^{2}\right)+8\left(1+4 \pi^{2} x^{2}\right)^{4} \\
& \left.-6\left(1+4 \pi^{2} x^{2}\right)^{3}\right\}\left(1+\pi^{2} x^{2}\right)^{7} H / \pi^{2} x^{2}\left(1+4 \pi^{2} x^{2}\right)^{3}\left(1+\pi^{2} x^{2}\right)^{7} . \tag{A.9}
\end{align*}
$$

To evaluate $\alpha_{r}{ }^{\prime \prime}$ we can transform the term $\delta^{\prime \prime}\left(\Delta E^{\prime}\right)$ in (2.6) as

$$
\begin{equation*}
\delta^{\prime \prime}\left(-\frac{\hbar k \hbar q}{m} v+\frac{\hbar^{2} q^{2}}{2 m}\right)=\frac{m^{3}}{\hbar^{3} k^{3} \hbar^{3} q^{3}} \delta^{\prime \prime}\left(v-\nu_{0}\right) \tag{A.10}
\end{equation*}
$$

The $d v$ integration is carried out going through the same steps as in (A.4) and (A.5), this time using the rule

$$
\begin{equation*}
f(x) \delta^{\prime \prime}(x)=f\left(x_{0}\right) \delta^{\prime \prime}(x)-2 f^{\prime}\left(x_{0}\right) \delta^{\prime}(x)+f^{\prime \prime}\left(x_{0}\right) \delta(x) \tag{A.11}
\end{equation*}
$$

The $\mathrm{d}(\hbar q)$ integral is lengthy but easy to carry out, and finally the $\mathrm{d}(\hbar k)$ integral results essentially in putting $k=k_{\mathrm{F}}$. We obtain

$$
\begin{align*}
\alpha_{r}^{\prime \prime}=\frac{V m^{3}}{3 N \hbar^{2} \pi} & \frac{4 \pi V}{(2 \pi)^{3} \hbar^{3}} \frac{\left(1+\pi^{2} x^{2}\right)^{4}}{384 \pi\left(1-A^{\prime}\right)}(64 \pi x)^{2}\left\{-960 \pi^{2} x^{2}\left(1+4 \pi^{2} x^{2}\right)^{4}\right. \\
& -72 \pi^{7} x^{9}\left(1+\pi^{2} x^{2}\right)^{8}+384\left(1+4 \pi^{2} x^{2}\right)^{4}+\left[\frac{27}{64}(x / \pi)-2.625 \pi x^{3}\right. \\
& \left.\left.-12.25 \pi^{3} x^{5}-7 \pi^{5} x^{7}\right]\left(1+\pi^{2} x^{2}\right)^{8}\right\} /\left(1+\pi^{2} x^{2}\right)^{8}\left(1+4 \pi^{2} x^{2}\right)^{4} . \tag{A.12}
\end{align*}
$$

In the mobility formula (1.1) $\alpha_{r}^{\prime \prime}$ appears in the numerator and $\alpha_{r}{ }^{\prime}$ in the denominator, and so to get the zero of $\mu$ we have to search where $\alpha_{r}{ }^{\prime \prime}$ goes zero. If we look at the curly bracket term in (A.12) we see that for $x=0$ it takes the value 384 ; as $x$ slowly increases it becomes small and passes through zero at $x=0.201290399297936$. After this value of $x$ the curly bracket term always remains negative and does not change sign. There is therefore no root other than $x=0.2$ in the region $0 \leqslant x \leqslant 1$. The whole expression is very sensitive to changes in the value of $x$; this is because of the high powers of $x$ appearing in the expression. Practically we may take the root to be $x=0.2$. Stated in another way this corresponds to $\rho_{A}^{1 / 3} a_{0}=0.2$.

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